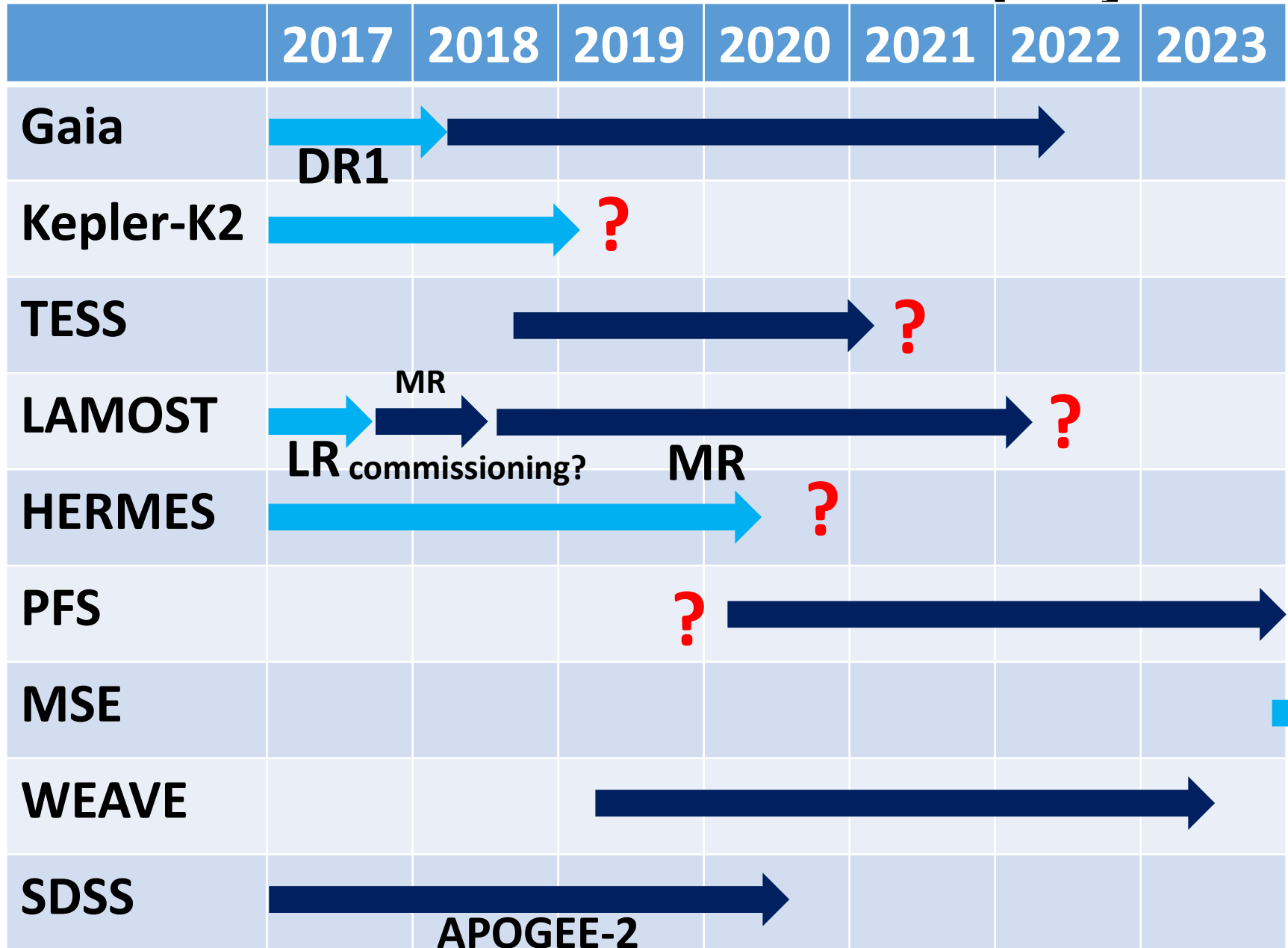


GA/Stars

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Timeline of instruments and projects



Which elements for chemical tagging?

...Having impact on choices of wavelength range and spectral resolution of new instruments.

- O, Mg: SN II ($M \geq 20M_{\text{sun}}$)
- Si, Ca, Ti, Cr (even-Z): SN II (not only massive ones)
- Mn, Co: SN II, metallicity dependent
- explosion energy dependent
- Eu: r-process. low-mass SN II?
- Y, Zr, Ba, La: s-process (+r-process at low metallicity)
- Fe, Ni: “metallicity”. SN Ia contribution
- determination of atmospheric parameters
- Sc, Y

Define some model cases of collaborations (both science and instrumentation)

- **Ongoing collaborations:**
LAMOST/Subaru-HDS for very metal-poor stars
- **Future follow-up**
Spectroscopic (HR) follow-up of objects from other surveys (e.g. PRISTINE, U-band imaging)
- **Instrumentation**
**Prototype of HR spectrometer for MSE on Subaru
(also prototype of Subaru/PFS-HR)**

Feedback to Subaru instrumentation and operation from scientific perspectives?

- **Faster response for follow-up?**

Launch science/instrumentation working group in each category for further discussion/contact/mini-workshop?

- **Discussion items:**
 - key elements for chemical tagging etc.
 - connecting spectrograph, spectral data ...
 - proposal for Subaru and other telescopes?
 - changing students
- **Meetings, email connection ...**
 - joining in meetings on PFS?